

Astrometric search for extrasolar planets

Tristan Röll

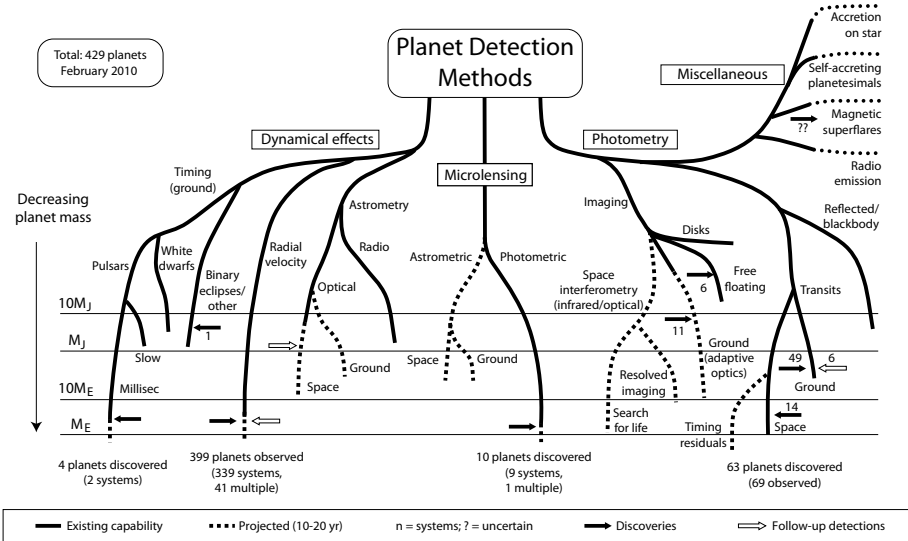
Ralph Neuhäuser, Andreas Seifahrt, Guillermo Torres,
Rainer Köhler, Jacob Bean

Astrophysikalisches Institut und Universitäts - Sternwarte Jena

Young Planetary Systems Workshop,
Jena 2010

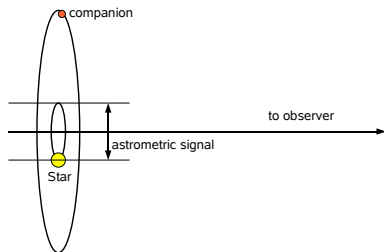


Exoplanet Detection - Methods



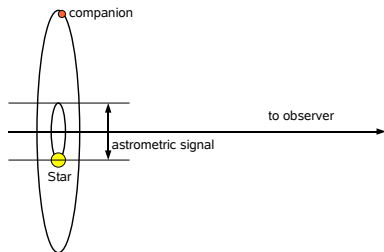
(diagram by M. Perryman)

Astrometric Signal



$$a_{\text{wobble}} [\text{AU}] = a_{\text{comp}} [\text{AU}] \frac{M_{\text{comp}}}{M_{\text{star}}}$$

Astrometric Signal

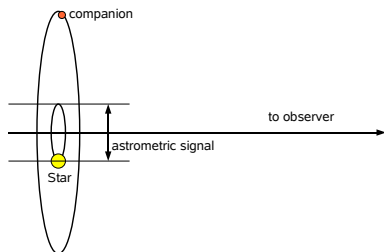


$$a_{\text{wobble}} [\text{AU}] = a_{\text{comp}} [\text{AU}] \frac{M_{\text{comp}}}{M_{\text{star}}}$$

Astrometric Signal (circular orbit)

$$\Theta [\text{mas}] = 2 \times a_{\text{wobble}} [\text{mas}] = 1.91 \times \frac{a_{\text{comp}} [\text{pc}]}{[\text{AU}]} \frac{M_{\text{comp}} [M_{\text{J}}]}{d} \frac{M_{\text{comp}} [M_{\odot}]}{M_{\text{star}}}$$

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Astrometric Signal (circular orbit)

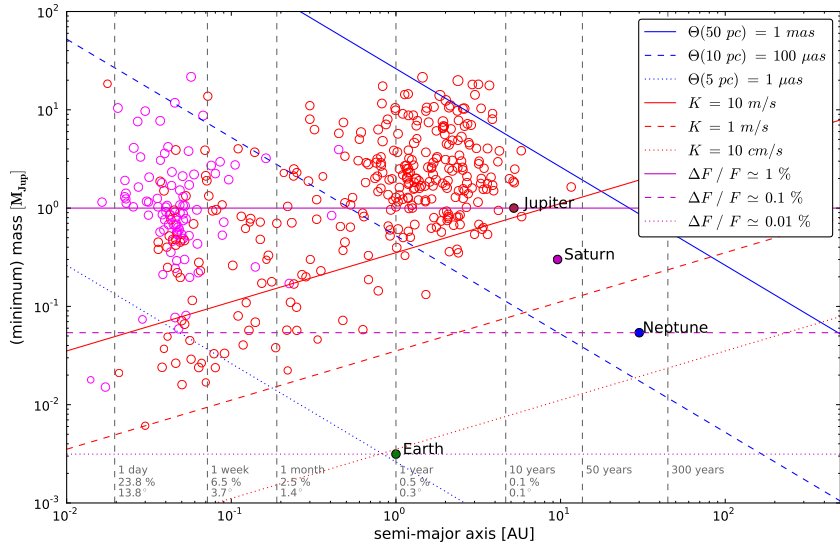
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Example

Jupiter with Sun seen from 5 pc / 100 pc : $\Theta = 2 \text{ mas} / 0.1 \text{ mas}$

Earth with Sun seen from 5 pc: $\Theta = 1.2 \mu\text{as}$

Exoplanet Detection - Sensitivity



(data from exoplanet.eu)

First Detection of an Astrometric Companion

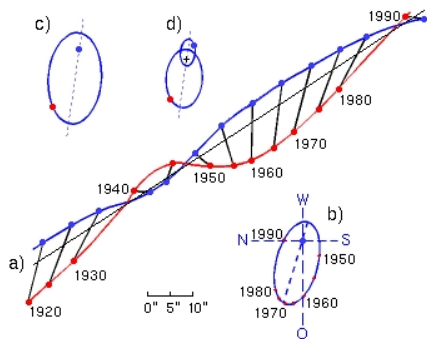
Friedrich Wilhelm Bessel (1784-1846)

1844

periodic perturbation in the proper motions of Sirius and Procyon

Conclusion

existence of unseen companion orbiting the star



Sirius A, Sirius B

Parallax, Proper Motion, Acceleration, and Orbital Motion of Barnard's Star

PETER VAN DE KAMP

Sproul Observatory, Swarthmore College, Swarthmore, Pennsylvania

(Received 31 December 1968)

Twenty-four early plates (1916–1919) and thirty consecutive years of photographic observations of Barnard's star covering the interval 1938–1967 confirm orbital motion with a period of 25 years and semi-axis major of $0''.0275$. The total number of plates is 3036, with 10 452 exposures, taken on 766 nights, with a total weight of 2056. Assuming a value of $0.15 \odot$ for the mass of Barnard's star, the mass of the companion is found to be $0.0016 \odot$, or 1.7 times the mass of Jupiter.

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- Sproul refractor (diameter: 61 cm, focal length: 10.9 m)
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- detection of systematic errors (Hershey, 1973)

Hubble Space Telescope

(launch: 1990)

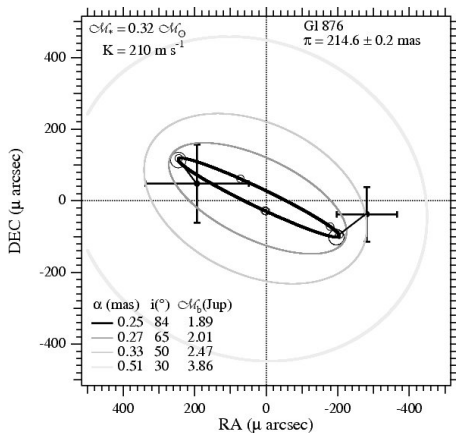
Currently the only telescope, which already delivers successful applications of astrometry regarding exoplanets.

Spaced based Astrometry

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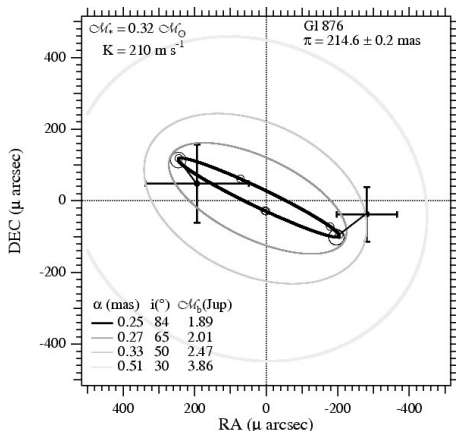
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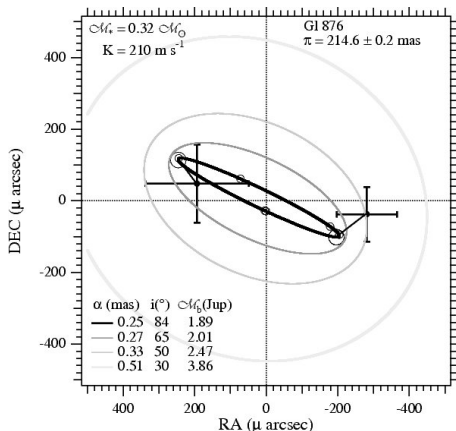
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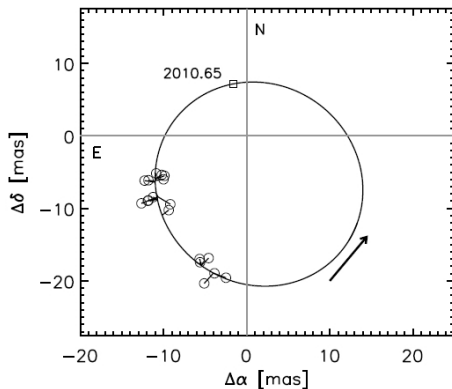
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- Gl 876 b
(Benedict et al., 2002)
- 55 Cnc d
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- Epsilon Eridani b
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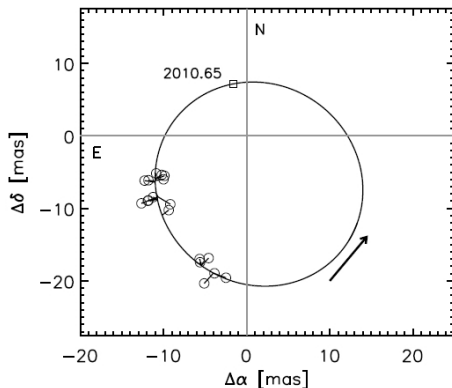
Spaced based Astrometry



HD 33636 b

RV planet candidate with
 $M \sin i = 9.3 M_J$

Spaced based Astrometry



HD 33636 b

RV planet candidate with
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Bean et al. 2007

determined with astrometry a true
mass of $M_{true} = 142 \pm 11 M_J$

Targets

nearby stellar multiple systems (distance < 100 pc)

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Telescope and Instruments

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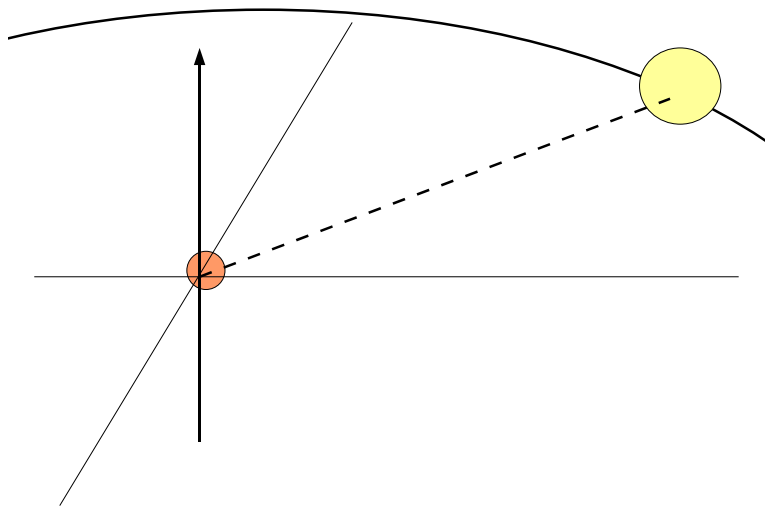
Telescope and Instruments

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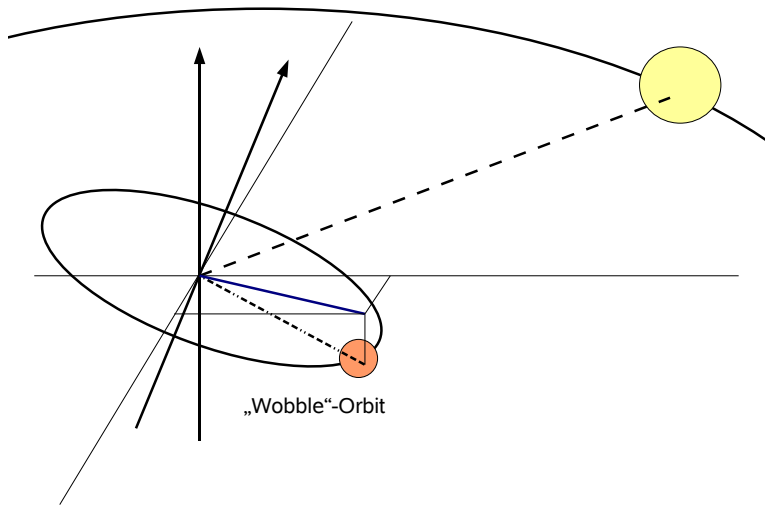
Calibration

old globular cluster (47Tuc and M15)

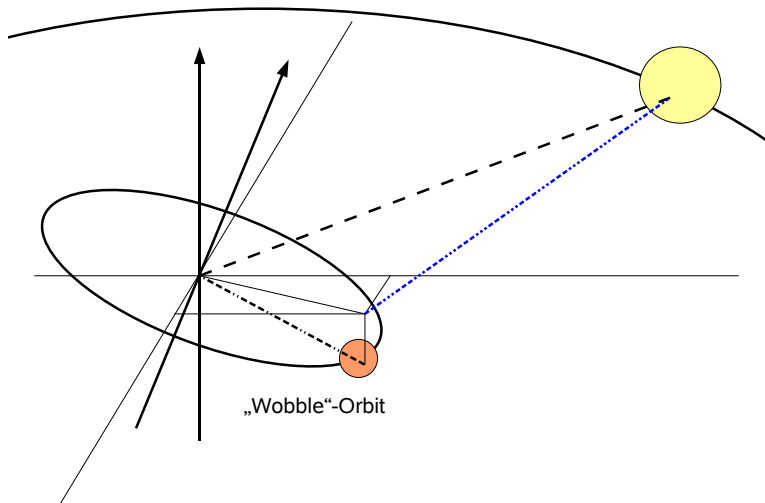
Astrometric Exoplanet Search at the AIU Jena



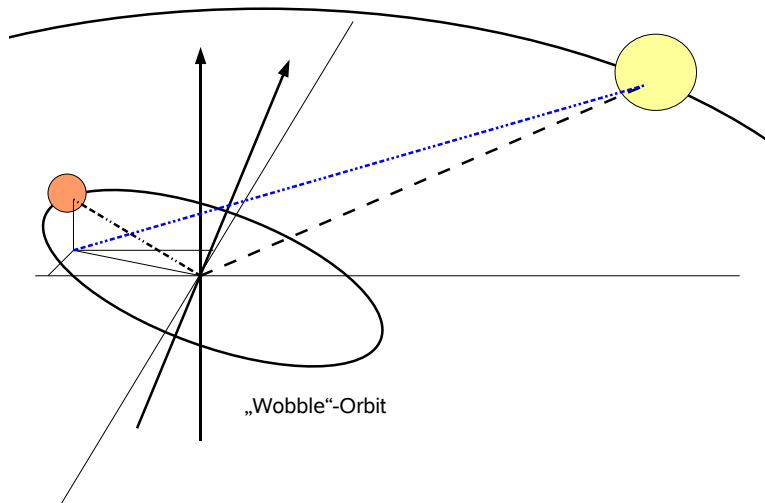
Astrometric Exoplanet Search at the AIU Jena



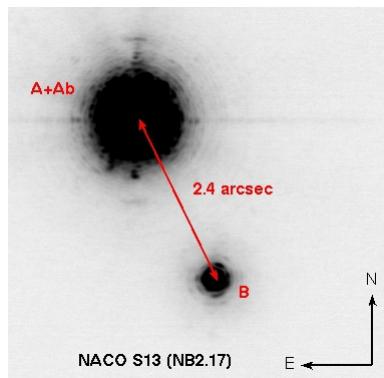
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First Target System: HD 19994 A&B



RV planet candidate HD 19994 Ab
Mayor et al. (2004)

$$P_{\text{pl}} \simeq 535 \text{ days}$$

$$a_{\text{pl}} \simeq 1.4 \text{ AU}$$

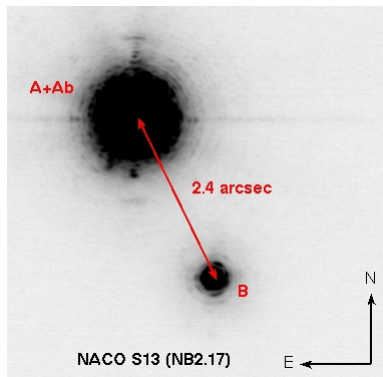
$$m \sin i \simeq 1.7 M_{\text{Jup}}$$

$$P_{\text{binary}} \geq 1400 \text{ years}$$

$$\text{distance} \simeq 22.6 \text{ pc}$$

$$M_{\text{A}} \simeq 1.34 M_{\odot}, \quad M_{\text{B}} \simeq 0.4 M_{\odot}$$

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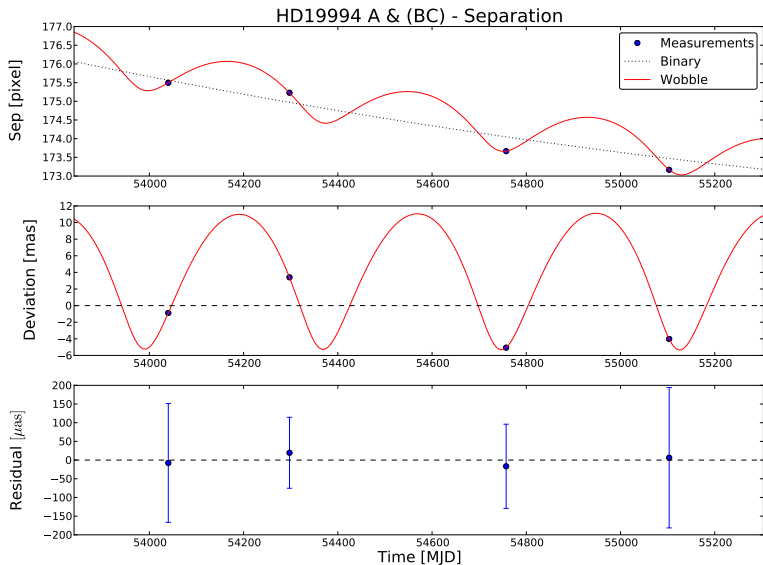
$$m \sin i \simeq 1.7 M_{\text{Jup}}$$

Expected Astrometric Signal:

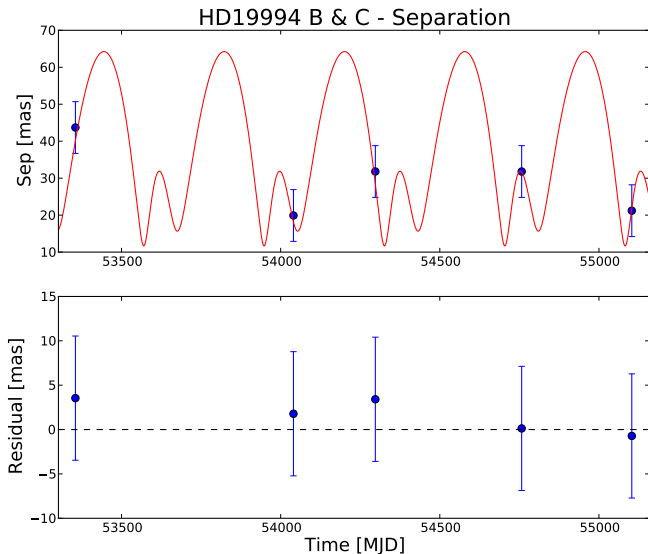
$$\Theta(i = 90^\circ) \simeq 150 \mu\text{as}$$

$$\Theta(13 M_{\text{Jup}}) \simeq 1.2 \text{ mas}$$

HD 19994 A&B - Astrometric measurements

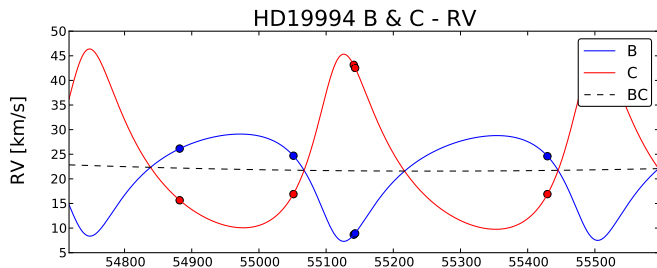


HD 19994 B & C - Speckle Interferometry



(speckle interferometry program written by Rainer Köhler)

HD 19994 B&C - Radial Velocity (CRIRES, VLT)



$$m_{\text{comp}} / M = 0.605$$

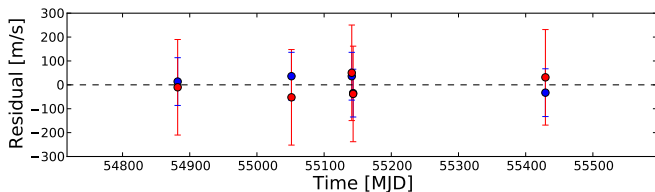
$$\text{Ecc} = 0.36$$

$$\text{Inc [deg]} = 108.63$$

$$P \text{ [days]} = 378.1$$

$$a_{\text{total}} \text{ [AU]} = 0.97$$

$$\omega \text{ [deg]} = 335.9$$

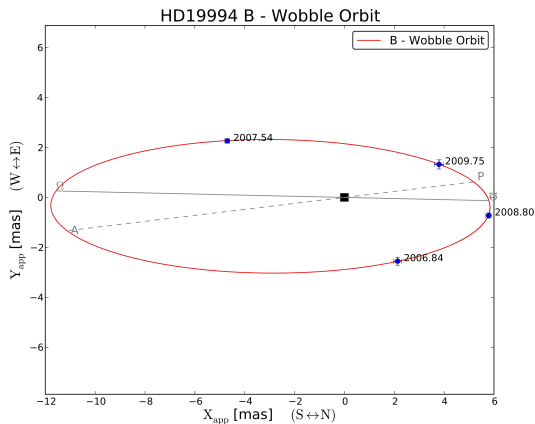


(radial velocity data by Andreas Seifahrt and Jacob Bean)



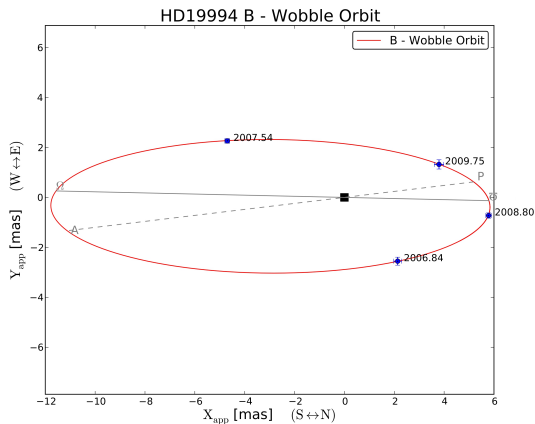
HD 19994 B - Wobble Orbit

$$\begin{aligned} P [\text{days}] &= 378.07^{+0.64}_{-0.60} \\ m_C [M_\odot] &= 0.326 \pm 0.012 \\ M_B [M_\odot] &= 0.54 \pm 0.014 \\ \text{Inc} [\text{deg}] &= 108.58^{+0.54}_{-0.51} \\ \text{Ecc} &= 0.36 \pm 0.014 \\ d [\text{pc}] &= 20.11^{+0.30}_{-0.27} \end{aligned}$$



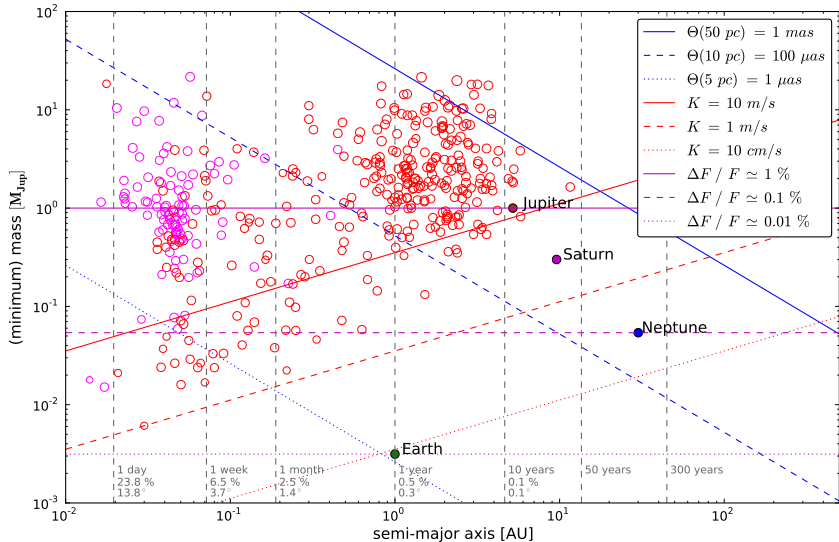
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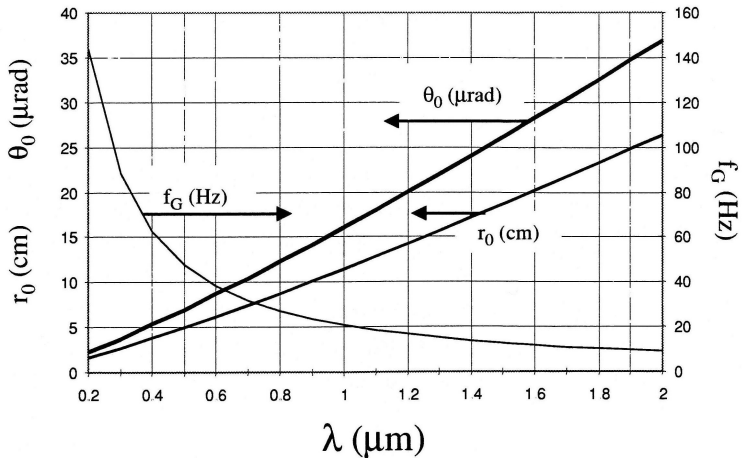
HD 19994 is a triple system, harboring one exoplanet

Exoplanet Detection - Sensitivity



(data from exoplanet.eu)

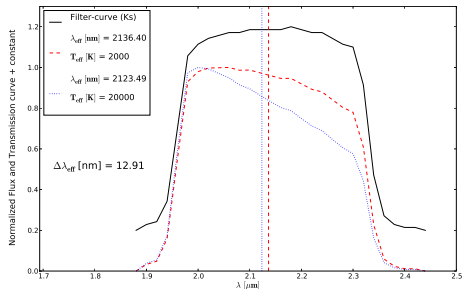
Earth's Atmospheric Turbulence's



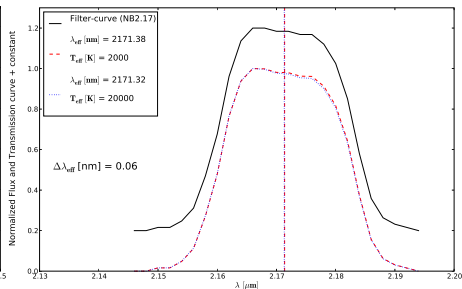
(by Tyson: "Introduction to Adaptive Optics")

Filter Transmission Curves

ESO Ks Filter



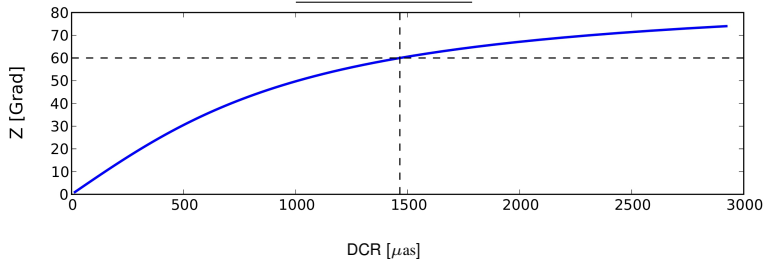
ESO NB2.17 Filter



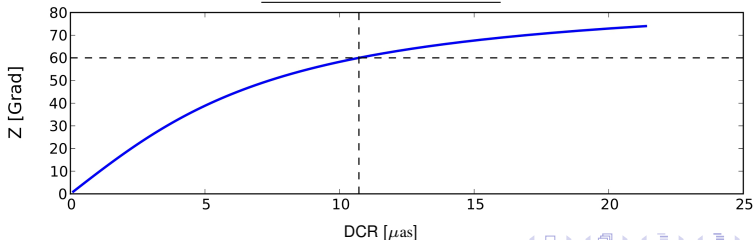
Extreme case: Two blackbodies with $T_{eff} = 2000 \text{ K}$ and $T_{eff} = 20000 \text{ K}$

Differential Chromatic Refraction

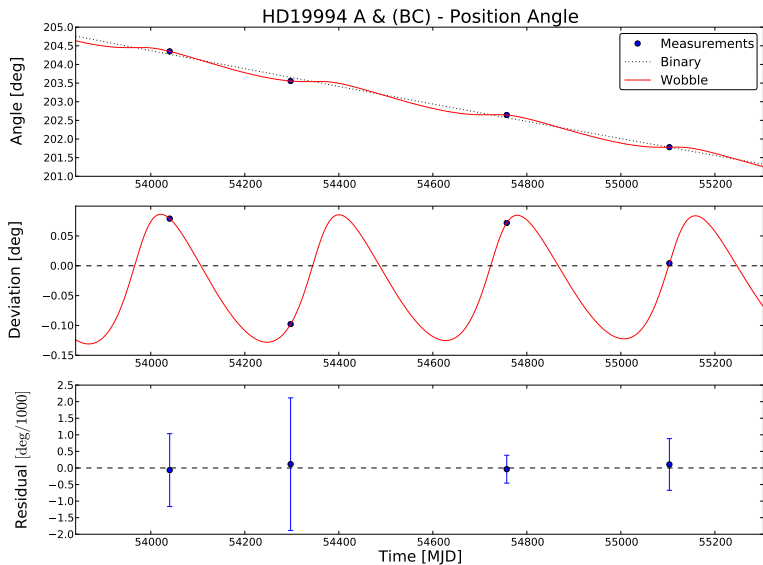
ESO Ks Filter



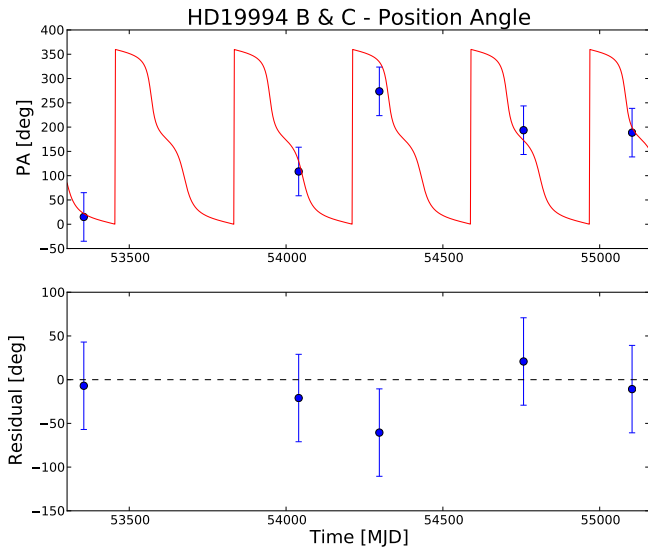
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HD 19994 A&B - Astrometric measurements



HD 19994 B&C - Speckle Interferometry



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